

Package ‘spatstat.univar’

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Title One-Dimensional Probability Distribution Support for the
'spatstat' Family

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Imports spatstat.utils (>= 3.2-1), graphics

Description Estimation of one-dimensional probability distributions
including kernel density estimation, weighted empirical cumulative
distribution functions, Kaplan-Meier and reduced-sample estimators
for right-censored data, heat kernels, kernel properties,
quantiles and integration.

License GPL (>= 2)

URL <http://spatstat.org/>

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BugReports <https://github.com/spatstat/spatstat.univar/issues>

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Contents

spatstat.univar-package	2
bw.abram	5
bw.abram.default	6
bw.pow	8
bw.taylor	9
CDF	11
densityAdaptiveKernel	12
densityAdaptiveKernel.default	12
densityBC	15
dkernel	19
dkernelBC	20
ewcdf	21
firstdigit	22
hotrod	24
indefinteg	25
integral	26
integral.density	27
kaplan.meier	28
kernel.factor	29
kernel.moment	30
kernel.squint	31
km.rs	32
knots.ewcdf	34
mean.ewcdf	35
quantile.density	36
quantile.ewcdf	37
quantilefun	38
reduced.sample	40
rounding	41
stiltjes	42
transformquantiles	44
uniquemap.default	45
unnormdensity	46
weighted.median	48
whist	50

Index

52

spatstat.univar-package
The spatstat.univar Package

Description

The **spatstat.univar** package belongs to the **spatstat** family of packages. It provides utilities for estimating the probability distribution of one-dimensional (real-valued) data.

Details

This package is a member of the **spatstat** family of packages. It provides utilities for estimation of the probability distribution of one-dimensional (i.e. numerical, real-valued) data. The utilities include:

- kernel density estimation:** including variable-bandwidth kernels, boundary correction, bandwidth selection, unnormalised weighted densities, and cumulative distribution functions of density estimates.
- weighted distributions and weighted statistics:** including weighted empirical cumulative distributions, weighted median, weighted quantiles, calculating the CDF from a density estimate
- estimation for right-censored data:** including Kaplan-Meier, reduced-sample and other estimators of the cumulative distribution function and hazard function from right-censored data
- quantiles:** including calculation of quantiles from an empirical cumulative distribution or a kernel density estimate
- kernels:** including calculation of the probability density, cumulative distribution function, quantiles, random generation, moments and partial moments of the standard smoothing kernels
- heat kernel:** calculation of the one-dimensional heat kernel in an interval
- integration:** Numerical integration including Stieltjes integrals and indefinite integrals.

The facilities are described in more detail below.

Kernel density estimation

The package supports fixed-bandwidth and variable-bandwidth kernel estimation of probability densities from numerical data. It provides boundary corrections for kernel estimates of densities on the positive half-line (applicable when the original observations are positive numbers) for both fixed-bandwidth and variable-bandwidth estimates.

If the observations have numerical weights associated with them, these weights will not be automatically normalised, and indeed the weights may be negative or zero. This is unlike the standard R method `density.default`.

The main functions are:

<code>unnormdensity</code>	extension of <code>density.default</code> allowing weights to be negative or zero.
<code>densityBC</code>	fixed-bandwidth kernel estimate with optional boundary correction
<code>densityAdaptiveKernel</code>	adaptive (variable-bandwidth) kernel estimation (generic)
<code>densityAdaptiveKernel.default</code>	adaptive (variable-bandwidth) kernel estimate (method for numeric data, with optional
<code>bw.abram.default</code>	calculate data-dependent bandwidths using Abramson rule
<code>CDF.density</code>	cumulative distribution function from kernel density estimate

Weighted distributions and weighted statistics

Weighted versions of standard operations such as the histogram and empirical distribution function are provided:

<code>whist</code>	weighted histogram
<code>ewcdf</code>	weighted empirical cumulative distribution function
<code>mean.ewcdf</code>	mean of weighted ecdf
<code>quantile.ewcdf</code>	quantiles of weighted ecdf

<code>knots.ewcdf</code>	jump points of weighted ecdf
<code>weighted.median</code>	weighted median of numeric values
<code>weighted.quantile</code>	weighted quantile of numeric values

Estimation for right-censored data

Facilities are provided for estimating the probability distribution of right-censored lifetimes (non-negative real random variables).

<code>kaplan.meier</code>	Kaplan-Meier estimator of cumulative distribution function and hazard rate, from right-censored data
<code>reduced.sample</code>	reduced-sample estimator of cumulative distribution function, from right-censored data

Quantiles

Facilities are provided for computing the quantiles of a probability distribution, given estimates of the probability density or the cumulative distribution function and so on.

<code>CDF.density</code>	cumulative distribution function from kernel density estimate
<code>quantile.density</code>	quantiles of kernel density estimate
<code>quantile.ewcdf</code>	quantiles of weighted ecdf
<code>quantilefun</code>	quantiles as a function
<code>quantilefun.ewcdf</code>	quantiles as a function
<code>weighted.quantile</code>	weighted quantile of numeric values
<code>transformquantiles</code>	transform the quantiles of a dataset

Kernels

The standard R function `density.default` recognises a list of smoothing kernels by name: "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". For these kernels, **spatstat.univar** provides various characteristics:

<code>dkernel</code>	probability density of the kernel
<code>pkernel</code>	cumulative distribution function of the kernel
<code>qkernel</code>	quantiles of the kernel
<code>rkernel</code>	generate simulated realisations from the kernel
<code>kernel.factor</code>	scale factor relating bandwidth to half-width of kernel
<code>kernel.moment</code>	partial moment of kernel
<code>kernel.squint</code>	integral of squared kernel
<code>dkernelBC</code>	evaluate the kernel with boundary correction

Heat kernels

The heat kernel in an interval can be calculated.

<code>hotrod</code>	calculate the heat kernel in an interval
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Integration

A few facilities are provided for calculating integrals of real functions.

<code>indefinteg</code>	indefinite integral
<code>integral.density</code>	integral of a kernel density estimate
<code>stieltjes</code>	Stieltjes integral

Utilities

A few utilities for numerical data are also provided.

<code>uniquemap.default</code>	map duplicates to unique entries
<code>rounding.default</code>	determine whether values have been rounded
<code>firstdigit</code>	leading digit in decimal representation
<code>lastdigit</code>	least significant digit in decimal representation
<code>ndigits</code>	number of digits in decimal representation

Licence

This library and its documentation are usable under the terms of the "GNU General Public License", a copy of which is distributed with the package.

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bw.abram

Abramson's Adaptive Bandwidths

Description

Computes adaptive smoothing bandwidths according to the inverse-square-root rule of Abramson (1982).

Usage

`bw.abram(X, h0, ...)`

Arguments

<code>X</code>	Data to be smoothed.
<code>h0</code>	Global smoothing bandwidth. A numeric value.
<code>...</code>	Additional arguments passed to methods.

Details

This function computes adaptive smoothing bandwidths for a dataset, using the methods of Abramson (1982) and Hall and Marron (1988).

The function `bw.abram` is generic. There is a default method `bw.abram.default`. The `spatstat` package family includes methods for spatial objects.

Value

See the documentation for the particular method.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

References

Abramson, I. (1982) On bandwidth variation in kernel estimates — a square root law. *Annals of Statistics*, **10**(4), 1217-1223.

Hall, P. and Marron, J.S. (1988) Variable window width kernel density estimates of probability densities. *Probability Theory and Related Fields*, **80**, 37-49.

Silverman, B.W. (1986) *Density Estimation for Statistics and Data Analysis*. Chapman and Hall, New York.

See Also

[bw.abram.default](#)

[bw.abram.default](#) *Abramson's Adaptive Bandwidths For Numeric Data*

Description

Computes adaptive smoothing bandwidths for numeric data, according to the inverse-square-root rule of Abramson (1982).

Usage

```
## Default S3 method:
bw.abram(X, h0, ...,
         at = c("data", "grid"),
         pilot = NULL, hp = h0, trim = 5, smoother = density.default)
```

Arguments

X Data for which bandwidths should be calculated. A numeric vector.

h0 A scalar value giving the global smoothing bandwidth in the same units as X. The default is $h0 = \text{bw.nrd0}(X)$.

... Arguments passed to [density.default](#) (or to smoother) controlling the range of values x at which the density will be estimated, when at="grid".

at Character string (partially matched) specifying whether to compute bandwidth values only at the data points of X (at = 'data', the default) or on a grid of x values (at = 'grid').

pilot	Optional. Specification of a pilot density (possibly unnormalised). Either a numeric vector giving the pilot density for each data point of X , a function in the R language, or a probability density estimate (object of class "density"). If <code>pilot=NULL</code> the pilot density is computed by applying fixed-bandwidth density estimation to X using bandwidth <code>hp</code> .
hp	Optional. A scalar pilot bandwidth, used for estimation of the pilot density, if <code>pilot</code> is not given.
trim	A trimming value required to curb excessively large bandwidths. See Details. The default is sensible in most cases.
smoother	Smoother for the pilot. A function or character string, specifying the function to be used to compute the pilot estimate when <code>pilot</code> is <code>NULL</code> .

Details

This function computes adaptive smoothing bandwidths using the methods of Abramson (1982) and Hall and Marron (1988).

The function `bw.abram` is generic. The function `bw.abram.default` documented here is the default method which is designed for numeric data.

If `at="data"` (the default) a smoothing bandwidth is computed for each data point in X . Alternatively if `at="grid"` a smoothing bandwidth is computed for a grid of x values.

Under the Abramson-Hall-Marron rule, the bandwidth at location u is

$$h(u) = h_0 * \min\left[\frac{\tilde{f}(u)^{-1/2}}{\gamma}, \text{trim}\right]$$

where $\tilde{f}(u)$ is a pilot estimate of the probability density. The variable bandwidths are rescaled by γ , the geometric mean of the $\tilde{f}(u)^{-1/2}$ terms evaluated at the data; this allows the global bandwidth h_0 to be considered on the same scale as a corresponding fixed bandwidth. The trimming value `trim` has the same interpretation as the required 'clipping' of the pilot density at some small nominal value (see Hall and Marron, 1988), to necessarily prevent extreme bandwidths (which can occur at very isolated observations).

The pilot density or intensity is determined as follows:

- If `pilot` is a function in the R language, this is taken as the pilot density.
- If `pilot` is a probability density estimate (object of class "density" produced by `density.default`) then this is taken as the pilot density.
- If `pilot` is `NULL`, then the pilot intensity is computed as a fixed-bandwidth kernel intensity estimate using `density.default` applied to the data X using the pilot bandwidth `hp`.

In each case the pilot density is renormalised to become a probability density, and then the Abramson rule is applied.

Instead of calculating the pilot as a fixed-bandwidth density estimate, the user can specify another density estimation procedure using the argument `smoother`. This should be either a function or the character string name of a function. It will replace `density.default` as the function used to calculate the pilot estimate. The pilot estimate will be computed as `smoother(X, sigma=hp, ...)` if `pilot` is `NULL`, or `smoother(pilot, sigma=hp, ...)` if `pilot` is a point pattern. If `smoother` does not recognise the argument name `sigma` for the smoothing bandwidth, then `hp` is effectively ignored.

Value

Either a numeric vector of the same length as *X* giving the Abramson bandwidth for each point (when *at* = "data", the default), or a function giving the Abramson bandwidths as a function of location.

Author(s)

Tilman Davies <Tilman.Davies@otago.ac.nz>. Adapted by Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

References

Abramson, I. (1982) On bandwidth variation in kernel estimates — a square root law. *Annals of Statistics*, **10**(4), 1217-1223.

Hall, P. and Marron, J.S. (1988) Variable window width kernel density estimates of probability densities. *Probability Theory and Related Fields*, **80**, 37-49.

Silverman, B.W. (1986) *Density Estimation for Statistics and Data Analysis*. Chapman and Hall, New York.

See Also

[bw.abram](#), [bw.nrd0](#).

Examples

```
xx <- rexp(20)
bw.abram(xx)
```

bw.pow

Variable Bandwidths Proportional to a Power of the Data Value

Description

Computes variable smoothing bandwidths intended to be proportional to the observed data values, raised to a given power.

Usage

```
bw.pow(X, h0, POW = 0.75, trim = 5, ...)
```

Arguments

X Data for which bandwidths should be calculated. A numeric vector of positive values.

h0 A scalar value giving the global smoothing bandwidth in the same units as *X*. The default is [bw.nrd0\(X\)](#).

POW Numeric value. The exponent of the power transformation to be applied to *X*.

trim	A trimming value required to curb excessively large bandwidths. See Details. The default is sensible in most cases.
...	Ignored.

Details

This function computes adaptive smoothing bandwidths for the data values in X . Larger data values are assigned larger bandwidths.

Bandwidths are proportional to X^{POW} . The bandwidth at location u is

$$h(u) = h_0 * \min\left[\frac{u^{POW}}{\gamma}, \text{trim}\right]$$

where γ is the geometric mean of the values u^{POW} . This allows the global bandwidth h_0 to be considered on the same scale as a corresponding fixed bandwidth.

Value

A numeric vector of the same length as X .

Author(s)

Tilman Davies <Tilman.Davies@otago.ac.nz>. Adapted by Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

See Also

[bw.abram](#), [bw.nrd0](#).

Examples

```
xx <- sort(rexp(10))
bb <- bw.pow(xx)
signif(rbind(xx, bb), 3)
```

bw.taylor

Bandwidth Selection for Kernel Density Estimation by Non-Random Bootstrap

Description

Use Taylor's non-random bootstrap technique to select the bandwidth for kernel density estimation on the real line.

Usage

```
bw.taylor(x, ..., strange = NULL, useC = TRUE)
```

Arguments

x	Numeric vector.
...	Ignored.
strange	Range of bandwidths to be considered. A numeric vector of length 2.
useC	Logical value specifying whether to use faster C code.

Details

This function selects a bandwidth for kernel density estimation of a probability density on the real line, using the numeric data `x` and assuming a Gaussian kernel. The result is the numeric value of the standard deviation of the Gaussian kernel.

The function uses the method of Taylor (1989) who showed that, when using the Gaussian kernel, the optimisation criterion can be computed rapidly from the data without any randomised resampling.

The domain of the probability density is assumed to be the entire real line. Boundary correction is not currently implemented.

The result of `bw.taylor` is a single numeric value giving the selected bandwidth.

Value

A single numeric value.

Author(s)

Tilman Davies <Tilman.Davies@otago.ac.nz> and Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

References

Taylor, C.C. (1989) Choice of the Smoothing Parameter in Kernel Density Estimation, *Biometrika* **76** 4, 705–712.

See Also

[bw.nrd](#) in the **stats** package for standard bandwidth selectors.

Examples

```
x <- rnorm(30)
bw.taylor(x)
```

CDF*Cumulative Distribution Function From Kernel Density Estimate*

Description

Given a kernel estimate of a probability density, compute the corresponding cumulative distribution function.

Usage

```
CDF(f, ...)

## S3 method for class 'density'
CDF(f, ..., warn = TRUE)
```

Arguments

<code>f</code>	Density estimate (object of class "density").
<code>...</code>	Ignored.
<code>warn</code>	Logical value indicating whether to issue a warning if the density estimate <code>f</code> had to be renormalised because it was computed in a restricted interval.

Details

`CDF` is generic, with a method for class "density".

This calculates the cumulative distribution function whose probability density has been estimated and stored in the object `f`. The object `f` must belong to the class "density", and would typically have been obtained from a call to the function [density](#).

Value

A function, which can be applied to any numeric value or vector of values.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>

See Also

[density](#), [quantile.density](#)

Examples

```
b <- density(runif(10))
f <- CDF(b)
f(0.5)
plot(f)
```

densityAdaptiveKernel *Adaptive Kernel Estimation of Density or Intensity*

Description

Computes an adaptive estimate of probability density or intensity using a variable-bandwidth smoothing kernel.

Usage

```
densityAdaptiveKernel(x, ...)
```

Arguments

x	Data to be smoothed.
...	Additional arguments passed to methods.

Details

This generic function computes an adaptive kernel estimate of probability density or intensity.

The function **densityAdaptiveKernel** is generic. The **spatstat** package family includes methods for spatial objects.

Value

See documentation for each method.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Tilman Davies <Tilman.Davies@otago.ac.nz>.

See Also

[bw.abram](#).

densityAdaptiveKernel.default*Adaptive Kernel Estimation of Probability Density*

Description

Computes an adaptive estimate of probability density from numeric data, using a variable-bandwidth smoothing kernel.

Usage

```
## Default S3 method:
densityAdaptiveKernel(X, bw, ...,
  weights = NULL,
  zerocor=c("none", "weighted", "convolution",
           "reflection", "bdrykern", "JonesFoster"),
  at = c("grid", "data"), ngroups=Inf, fast=TRUE)
```

Arguments

X	Data to be smoothed. A numeric vector.
bw	Smoothing bandwidths. Either a numeric vector of the same length as X giving the bandwidth associated with each data value, or a function in the R language that provides the smoothing bandwidth at any desired location. The default is to compute bandwidths using <code>bw.abram.default</code> .
...	Additional arguments passed to <code>density.default</code> controlling the range of x values at which the density must be estimated, when <code>at="grid"</code> .
weights	Optional. Numeric vector of weights attached to each value in X.
zerocor	Character string (partially matched) specifying a boundary correction. This is appropriate when X contains only positive values.
at	String (partially matched) specifying whether to evaluate the probability density only at the data points (<code>at="data"</code>) or on a grid of x values (<code>at="grid"</code> , the default).
ngroups	Integer, Inf or NULL. If <code>ngroups = Inf</code> , the density estimate will be computed exactly using C code. If <code>ngroups</code> is finite, then the fast subdivision technique of Davies and Baddeley (2018) will be applied. If <code>ngroups = NULL</code> then a default rule is used to choose an efficient number of groups.
fast	Logical value specifying whether to use the Fast Fourier Transform to accelerate computations, when appropriate.

Details

This function computes an adaptive kernel estimate of probability density on the real line (if `zerocor="none"` or on the positive real line (if `zerocor` is another value).

The argument `bw` specifies the smoothing bandwidths to be applied to each of the points in `X`. It may be a numeric vector of bandwidth values, or a function yielding the bandwidth values.

If the values in `X` are x_1, \dots, x_n and the corresponding bandwidths are $\sigma_1, \dots, \sigma_n$ then the adaptive kernel estimate of intensity at a location u is

$$\hat{\lambda}(u) = \sum_{i=1}^n k(u, x_i, \sigma_i)$$

where $k(u, v, \sigma)$ is the value at u of the (possibly edge-corrected) smoothing kernel with bandwidth σ induced by a data point at v .

Exact computation of the estimate above can be time-consuming: it takes n times longer than fixed-bandwidth smoothing.

The partitioning method of Davies and Baddeley (2018) accelerates this computation by partitioning the range of bandwidths into `ngroups` intervals, correspondingly subdividing `X` into `ngroups` subsets according to `bandwidth`, and applying fixed-bandwidth smoothing to each subset.

If `ngroups=NULL` then we use a default rule where `ngroups` is the integer part of the square root of the number of points in `X`, so that the computation time is only about \sqrt{n} times slower than fixed-bandwidth smoothing. Any positive value of `ngroups` can be specified by the user. Specifying `ngroups=Inf` enforces exact computation of the estimate without partitioning. Specifying `ngroups=1` is the same as fixed-bandwidth smoothing with `bandwidth sigma=median(bw)`.

Value

If `at="data"`, a numeric vector of the same length as `X`. If `at="grid"`, a probability density object of class "density".

Bandwidths and Bandwidth Selection

The function `densityAdaptiveKernel.default` computes one adaptive estimate of probability density, determined by the smoothing bandwidth values `bw`.

Typically the bandwidth values are computed by first computing a pilot estimate of the intensity, then using `bw.abram.default` to compute the vector of bandwidths according to Abramson's rule. This involves specifying a global bandwidth `h0`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Tilman Davies <Tilman.Davies@otago.ac.nz>.

References

Davies, T.M. and Baddeley, A. (2018) Fast computation of spatially adaptive kernel estimates. *Statistics and Computing*, **28**(4), 937-956.

Hall, P. and Marron, J.S. (1988) Variable window width kernel density estimates of probability densities. *Probability Theory and Related Fields*, **80**, 37-49.

Silverman, B.W. (1986) *Density Estimation for Statistics and Data Analysis*. Chapman and Hall, New York.

See Also

`bw.abram.default`

Examples

```
xx <- rexp(100, rate=5)
plot(density(xx))
curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid"))
curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid", zerocor="w"))
```

```

curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid", zerocor="c"))
curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid", zerocor="r"))
curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid", zerocor="b"))
curve(5 * exp(-5 * x), add=TRUE, col=3)
plot(densityAdaptiveKernel(xx, at="grid", zerocor="J"))
curve(5 * exp(-5 * x), add=TRUE, col=3)

```

densityBC

Kernel Density Estimation with Optional Boundary Correction

Description

Fixed-bandwidth kernel density estimation on the real line, or the positive real half-line, including optional corrections for a boundary at zero.

Usage

```

densityBC(x, kernel = "epanechnikov", bw=NULL,
           ...,
           h=NULL,
           adjust = 1,
           weights = rep(1, length(x))/length(x), from, to = max(x), n = 256,
           zerocor = c("none", "weighted", "convolution", "reflection",
                      "bdrykern", "JonesFoster"),
           fast=FALSE,
           internal=list())

```

Arguments

<code>x</code>	Numeric vector of observed values.
<code>kernel</code>	String specifying kernel. Options are "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". (Partial matching is used). Options are described in the help for density.default .
<code>bw, h</code>	Alternative specifications of the scale factor for the kernel. The bandwidth <code>bw</code> is the standard deviation of the kernel (this agrees with the argument <code>bw</code> in density.default). The rescale factor <code>h</code> is the factor by which the 'standard form' of the kernel is rescaled. For the Epanechnikov kernel, <code>h = bw * sqrt(5)</code> is the half-width of the support, while for the Gaussian kernel, <code>h = bw</code> is the standard deviation. Either <code>bw</code> or <code>h</code> should be given, and should be a single numeric value, or a character string indicating a bandwidth selection rule as described in density.default .
<code>adjust</code>	Numeric value used to rescale the bandwidth <code>bw</code> and halfwidth <code>h</code> . The bandwidth used is <code>adjust * bw</code> . This makes it easy to specify values like 'half the default' bandwidth.

weights	Numeric vector of weights associated with x. The weights are not required to sum to 1, and will not be normalised to sum to 1. The weights may include negative values.
from, to	Lower and upper limits of interval on which density should be computed. The default value of from is from=min(x) if zeroCor="none", and from=0 otherwise.
n	Number of r values for which density should be computed.
zeroCor	String (partially matched) specifying a correction for the boundary effect bias at $r = 0$ when estimating a density on the positive half line. Possible values are "none", "weighted", "convolution", "reflection", "bdrykern" and "JonesFoster".
fast	Logical value specifying whether to perform the calculation rapidly using the Fast Fourier Transform (fast=TRUE) or to use slower, exact code (fast=FALSE, the default).
internal	Internal use only.
...	Additional arguments are ignored.

Details

This function computes a fixed-bandwidth kernel estimate of a probability density on the real line, or the positive half-line, including optional boundary corrections for truncation of the density onto the positive half line.

Weighted estimates are supported, including negative weights. Weights are not renormalised to sum to 1. The resulting probability density estimate is not renormalised to integrate to 1.

Options for the smoothing kernel are described in the help for [density.default](#). The default is the Epanechnikov (truncated quadratic) kernel.

If zeroCor is missing, or given as "none", this function computes the fixed-bandwidth kernel estimator of the probability density on the real line, using [density.default](#). The estimated probability density (unnormalised) is

$$\hat{f}(x) = \sum_{i=1}^n w_i \kappa(x - x_i)$$

where x_1, \dots, x_n are the data values, w_1, \dots, w_n are the weights (defaulting to $w_i = 1/n$), and κ is the kernel, a probability density on the real line.

If zeroCor is given, the probability density is assumed to be confined to the positive half-line; the numerical values in x must all be non-negative; and a boundary correction is applied to compensate for bias arising due to truncation at the origin:

zeroCor="weighted": The contribution from each data point x_i is weighted by the factor $1/m(x_i)$ where $m(x) = 1 - F(-x)$ is the total mass of the kernel centred on x that lies in the positive half-line, and $F(x)$ is the cumulative distribution function of the kernel. The corrected estimate is

$$\hat{f}_W(x) = \sum_{i=1}^n w_i \frac{\kappa(x - x_i)}{1 - F(-x_i)}$$

This is the "cut-and-normalization" method of Gasser and Müller (1979). Effectively the kernel is renormalized so that it integrates to 1, and the adjusted kernel conserves mass.

zerocor="convolution": The estimate of the density $f(x)$ is weighted by the factor $1/m(x)$ where $m(r) = 1 - F(-x)$ is given above. The corrected estimate is

$$\hat{f}_C(x) = \sum_{i=1}^n w_i \frac{\kappa(x - x_i)}{1 - F(-x)}$$

This is the “convolution”, “uniform” or “zero-order” boundary correction method often attributed to Diggle (1985). This correction does not conserve mass. It is faster to compute than the weighted correction.

zerocor="reflection": if the kernel centred at data point x_i has a tail that lies on the negative half-line, this tail is reflected onto the positive half-line. The corrected estimate is

$$\hat{f}_R(x) = \sum_{i=1}^n w_i [\kappa(x - x_i) + \kappa(-x - x_i)]$$

This is the “reflection” method first proposed by Boneva et al (1971). This correction conserves mass. The estimated density always has zero derivative at the origin, $\hat{f}'_R(0) = 0$, which may or may not be desirable.

zerocor="bdrykern": The density estimate is computed using the Linear Boundary Kernel associated with the chosen kernel (Wand and Jones, 1995, page 47). The estimated (unnormalised) probability density is

$$\hat{f}_B(x) = \sum_{i=1}^n w_i [A(x) + (x - x_i)B(x)]\kappa(x - x_i)$$

where $A(x) = a_2(x)/D(x)$ and $B(x) = -a_1(x)/D(x)$ with $D(x) = a_0(x)a_2(x) - a_1(x)^2$ where $a_k(x) = \int_{-\infty}^x t^k \kappa(t) dt$. That is, when estimating the density $f(x)$ for values of x close to zero (defined as $x < h$ for all kernels except the Gaussian), the kernel contribution $k_h(x - x_i)$ is multiplied by a term that is a linear function of $x - x_i$, with coefficients depending on x . This correction does not conserve mass and may result in negative values, but is asymptotically optimal. Computation time for this estimate is greater than for the options above.

zerocor="JonesFoster": The modification of the Boundary Kernel estimate proposed by Jones and Foster (1996) is computed:

$$\hat{f}_{JF}(x) = \hat{f}_C(x) \exp\left(\frac{\hat{f}_B(x)}{\hat{f}_C(r)} - 1\right)$$

where $\hat{f}_C(r)$ is the convolution estimator and $\hat{f}_B(r)$ is the linear boundary kernel estimator. This ensures that the estimate is always nonnegative and retains the asymptotic optimality of the linear boundary kernel. Computation time for this estimate is greater than for all the options above.

If `fast=TRUE`, the calculations are performed rapidly using `density.default` which employs the Fast Fourier Transform. If `fast=FALSE` (the default), the calculations are performed exactly using slower C code.

Value

An object of class "density" as described in the help file for [density.default](#). It contains at least the entries

x	Vector of x values
y	Vector of density values $y = f(x)$

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Martin Hazelton <Martin.Hazelton@otago.ac.nz>.

References

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Gasser, Th. and Müller, H.-G. (1979). Kernel estimation of regression functions. In Th. Gasser and M. Rosenblatt (editors) *Smoothing Techniques for Curve Estimation*, pages 23–68. Springer, Berlin.

Jones, M.C. and Foster, P.J. (1996) A simple nonnegative boundary correction method for kernel density estimation. *Statistica Sinica*, **6** (4) 1005–1013.

Wand, M.P. and Jones, M.C. (1995) *Kernel Smoothing*. Chapman and Hall.

See Also

[density.default](#).

[dkernel](#) for the kernel itself.

[densityAdaptiveKernel.default](#) for adaptive (variable-bandwidth) estimation.

Examples

```
sim.dat <- rexp(500)
fhatN <- densityBC(sim.dat, "biweight", h=0.4)
fhatB <- densityBC(sim.dat, "biweight", h=0.4, zerocor="bdrykern")
plot(fhatN, ylim=c(0,1.1), main="density estimates")
lines(fhatB, col=2)
curve(dexp(x), add=TRUE, from=0, col=3)
legend(2, 0.8,
       legend=c("fixed bandwidth", "boundary kernel", "true density"),
       col=1:3, lty=rep(1,3))
```

Description

Density, distribution function, quantile function and random generation for several distributions used in kernel estimation for numerical data.

Usage

```
dkernel(x, kernel = "gaussian", mean = 0, sd = 1)
pkernel(q, kernel = "gaussian", mean = 0, sd = 1, lower.tail = TRUE)
qkernel(p, kernel = "gaussian", mean = 0, sd = 1, lower.tail = TRUE)
rkernel(n, kernel = "gaussian", mean = 0, sd = 1)
```

Arguments

x, q	Vector of quantiles.
p	Vector of probabilities.
kernel	String name of the kernel. Options are "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". (Partial matching is used).
n	Number of observations.
mean	Mean of distribution.
sd	Standard deviation of distribution.
lower.tail	logical; if TRUE (the default), then probabilities are $P(X \leq x)$, otherwise, $P(X > x)$.

Details

These functions give the probability density, cumulative distribution function, quantile function and random generation for several distributions used in kernel estimation for one-dimensional (numerical) data.

The available kernels are those used in [density.default](#), namely "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". For more information about these kernels, see [density.default](#).

dkernel gives the probability density, pkernel gives the cumulative distribution function, qkernel gives the quantile function, and rkernel generates random deviates.

Value

A numeric vector. For dkernel, a vector of the same length as x containing the corresponding values of the probability density. For pkernel, a vector of the same length as x containing the corresponding values of the cumulative distribution function. For qkernel, a vector of the same length as p containing the corresponding quantiles. For rkernel, a vector of length n containing randomly generated values.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Martin Hazelton <Martin.Hazelton@otago.ac.nz>.

See Also

[density.default](#), [kernel.factor](#), [kernel.moment](#), [kernel.squint](#).

Examples

```
x <- seq(-3,3,length=100)
plot(x, dkernel(x, "epa"), type="l",
      main=c("Epanechnikov kernel", "probability density"))
plot(x, pkernel(x, "opt"), type="l",
      main=c("OptCosine kernel", "cumulative distribution function"))
p <- seq(0,1, length=256)
plot(p, qkernel(p, "biw"), type="l",
      main=c("Biweight kernel", "cumulative distribution function"))
y <- rkernel(100, "tri")
hist(y, main="Random variates from triangular density")
rug(y)
```

Description

Computes the boundary-corrected version of a smoothing kernel density function.

Usage

```
dkernelBC(x, mean, sd = 1, kernel = "gaussian",
          zerocor = c("none", "weighted", "convolution",
                     "reflection", "bdrykern"))
```

Arguments

<code>x</code>	Numeric. Values of the function argument, at which the function should be evaluated.
<code>mean</code>	Numeric. The mean of the uncorrected kernel.
<code>sd</code>	Numeric value. The standard deviation of the uncorrected kernel.
<code>kernel</code>	Character string giving the name of the kernel as recognised by match.kernel .
<code>zerocor</code>	String (partially matched) specifying a correction for the boundary effect bias at $r = 0$ when estimating a density on the positive half line. Possible values are "none", "weighted", "convolution", "reflection", and "bdrykern".

Details

The kernel density function identified by `kernel` with standard deviation `sd` and mean `mean` will be computed, and truncated onto the positive half-line. The boundary correction specified by `zerocor` will then be applied. The result is the vector of corrected density values.

Value

Numeric value or numeric vector.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

See Also

[densityBC](#) to compute a density estimate using the boundary-corrected kernel.
[dkernel](#) to compute the un-corrected kernel density function, and [density.default](#) to compute an uncorrected density estimate.
[match.kernel](#) for the list of recognised names of kernels.

Examples

```
curve(dkernelBC(x, mean=1, zerocor="none"), to=5)
curve(dkernelBC(x, mean=1, zerocor="weighted"), to=5)
curve(dkernelBC(x, mean=1, zerocor="reflection"), to=5)
curve(dkernelBC(x, mean=1, zerocor="convolution"), to=5)
curve(dkernelBC(x, mean=1, zerocor="bdrykern"), to=5)
```

ewcdf

Weighted Empirical Cumulative Distribution Function

Description

Compute a weighted version of the empirical cumulative distribution function.

Usage

```
ewcdf(x, weights = NULL, normalise=TRUE, adjust=1)
```

Arguments

<code>x</code>	Numeric vector of observations.
<code>weights</code>	Optional. Numeric vector of non-negative weights for <code>x</code> . Defaults to equal weight 1 for each entry of <code>x</code> .
<code>normalise</code>	Logical value indicating whether the weights should be rescaled so that they sum to 1.
<code>adjust</code>	Numeric value. Adjustment factor. The weights will be multiplied by <code>adjust</code> .

Details

This is a modification of the standard function [ecdf](#) allowing the observations x to have weights.

The weighted e.c.d.f. (empirical cumulative distribution function) F_n is defined so that, for any real number y , the value of $F_n(y)$ is equal to the total weight of all entries of x that are less than or equal to y . That is $F_n(y) = \text{sum}(\text{weights}[x \leq y])$.

Thus F_n is a step function which jumps at the values of x . The height of the jump at a point y is the total weight of all entries in x number of tied observations at that value. Missing values are ignored.

If `weights` is omitted, the default is equivalent to `ecdf(x)` except for the class membership.

The result of `ewcdf` is a function, of class "ewcdf", inheriting from the classes "ecdf" (only if `normalise=TRUE`) and "stepfun".

The class `ewcdf` has methods for [print](#), [quantile](#) and [mean](#).

The inherited classes `ecdf` and `stepfun` have methods for [plot](#) and [summary](#).

Value

A function, of class "ewcdf", inheriting from "ecdf" (if `normalise=TRUE`) and "stepfun".

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

[ecdf](#).

[quantile.ewcdf](#), [mean.ewcdf](#).

Integrals with respect to the weighted cumulative distribution function can be computed using [stieljes](#).

Examples

```
x <- rnorm(100)
w <- runif(100)
plot(e <- ewcdf(x,w))
e
```

Description

Find the first or last digit in the decimal representation of a number.

Usage

```
firstdigit(x)
lastdigit(x)
ndigits(x)
```

Arguments

`x` A numeric value or numeric vector.

Details

`firstdigit(x)` finds the first (most significant) digit, `lastdigit(x)` finds the last (least significant) digit, and `ndigits(x)` finds the number of digits, in the decimal representation of each entry of `x`. The decimal representation is truncated at the number of digits available for double precision numbers on the hardware, usually 15.

Value

An integer or integer vector of the same length as `x`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>

See Also

[rounding](#)

Examples

```
firstdigit(42)
lastdigit(42)
ndigits(42)

firstdigit(-0.1234)
lastdigit(-0.1234)
ndigits(-0.1234)

firstdigit(0)
lastdigit(0)
ndigits(0)
```

hotrod

*Heat Kernel for a One-Dimensional Rod***Description**

Calculate values of the heat kernel on a one-dimensional rod. The ends of the rod may be assumed to be insulated, or absorbing.

Usage

```
hotrod(len, xsource, xquery, sigma, ends=c("insulated", "absorbing"), nmax=20)
```

Arguments

len	Length of the rod. A single number or numeric vector.
xsource	Positions of the source points, from the left end of the rod (in the same distance units as len). A single number or numeric vector.
xquery	Positions of the query points, from the left end of the rod (in the same distance units as len). A single number or numeric vector.
sigma	Bandwidth for kernel. A single number or a numeric vector.
ends	Character string (partially matched) specifying whether the ends of the rod are assumed to be insulated or absorbing.
nmax	Number of terms in the infinite sum to use. A single integer or an integer vector.

Details

Computes the heat kernel as an infinite sum.

Value

Number or numeric vector.

Author(s)

Greg McSwiggan and Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

Examples

```
curve(hotrod(1, 0.1, x, 0.7))

# check it's a probability density
f <- function(x) hotrod(1, 0.1, x, 0.7)
integrate(f, 0, 1)

## absorbing ends
curve(hotrod(1, 0.1, x, 0.7, ends="a"))
```

indefinteg*Indefinite Integral*

Description

Computes the indefinite integral of the given function.

Usage

```
indefinteg(f, x, ...,
           method=c("trapezoid", "quadrature"),
           lower=min(x), nfine=8192)
```

Arguments

<code>f</code>	an R function taking a numeric first argument and returning a numeric vector of the same length.
<code>x</code>	Vector of values of the argument for which the indefinite integral should be evaluated.
<code>...</code>	additional arguments to be passed to <code>f</code> .
<code>method</code>	String (partially matched) specifying how to compute the integrals.
<code>lower</code>	Lower limit of integration. A single number.
<code>nfine</code>	Number of sub-intervals to use for computation if <code>method='trapezoid'</code> .

Details

The indefinite integral of the given function `f` is computed numerically at each of the desired values `x`. The lower limit of integration is taken to be `min(x)`.

The result is a numeric vector `y` of the same length as `x`, with entries

$$y_i = \int_{\text{lower}}^{x_i} f(t) dt$$

If `method='trapezoid'` (the default), the integrals are computed rapidly using the trapezoid rule. If `method='quadrature'` the integrals are computed accurately but much more slowly, using the numerical quadrature routine [integrate](#).

If `method='trapezoid'` the function `f` is first evaluated on a finer grid of values of the function argument. The fine grid contains `nfine` sample points. The values of the indefinite integral on the fine grid are computed using the trapezoidal approximation. Finally the values of the indefinite integral are extracted at the desired argument values `x`.

Value

Numeric vector of the same length as `x`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

See Also

[integrate](#)

Examples

```
curve(indefinteg(sin, x), to=pi)
```

integral

Integral of a Function or Spatial Object

Description

Computes the integral of a function or spatial object.

Usage

```
integral(f, domain=NULL, ...)
```

Arguments

- `f` A function, or a spatial object that can be treated as a function.
- `domain` Optional. Data specifying the domain of integration.
- `...` Arguments passed to methods.

Details

The function `integral` is generic. It calculates the integral of a function, or the integral of a spatial object that can be treated as a function. It has methods for one-dimensional functions ("density", "fv") and for spatial objects ("im", "msr", "linim", "linfun").

Value

A single numeric or complex value, or a vector of such values.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

[integral.density](#).

`integral.im` in package `spatstat.geom`.

<code>integral.density</code>	<i>Compute Integral of One-Dimensional Kernel Density Estimate.</i>
-------------------------------	---

Description

Compute the integral of a kernel density estimate over a specified range.

Usage

```
## S3 method for class 'density'
integral(f, domain = NULL, weight=NULL, ...)
```

Arguments

<code>f</code>	A one-dimensional probability density estimate (object of class "density") obtained from the function <code>density.default</code> or from <code>unnormdensity</code> .
<code>domain</code>	Optional. Range of values of the argument x over which the density $f(x)$ should be integrated. A numeric vector of length 2 giving the minimum and maximum values of x . Infinite limits are permitted.
<code>weight</code>	Optional. A function(x) specifying a weight integrand.
...	Ignored.

Details

This is a method for the generic function `integral`. It computes the numerical integral

$$I = \int f(x)dx$$

of the density estimate `f`. If `weight` is specified, then the weighted integral

$$I = \int w(x)f(x)dx$$

is computed, where `w` is the function specified by `weight`. This function must return finite numerical values.

If `domain` is specified, the integral is restricted to the interval of x values given by the `domain`.

Integrals are calculated numerically using the trapezoidal rule restricted to the `domain` given.

Value

A single numerical value.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

See Also

[density.default](#)
[quantile.density, CDF.density](#)

Examples

```
x <- runif(10)
d <- density(x, bw=0.1)
integral(d) # should be approximately 1
integral(d, domain=c(-Inf, 0)) # mass on negative half-line
## mean of density
integral(d, weight=function(x) x)
```

kaplan.meier

*Kaplan-Meier Estimator using Histogram Data***Description**

Compute the Kaplan-Meier estimator of a survival time distribution function, from histogram data

Usage

```
kaplan.meier(obs, nco, breaks, upperobs=0)
```

Arguments

obs	vector of n integers giving the histogram of all observations (censored or uncensored survival times)
nco	vector of n integers giving the histogram of uncensored observations (those survival times that are less than or equal to the censoring time)
breaks	Vector of $n + 1$ breakpoints which were used to form both histograms.
upperobs	Number of observations beyond the rightmost breakpoint, if any.

Details

This function is needed mainly for internal use in **spatstat**, but may be useful in other applications where you want to form the Kaplan-Meier estimator from a huge dataset.

Suppose T_i are the survival times of individuals $i = 1, \dots, M$ with unknown distribution function $F(t)$ which we wish to estimate. Suppose these times are right-censored by random censoring times C_i . Thus the observations consist of right-censored survival times $\tilde{T}_i = \min(T_i, C_i)$ and non-censoring indicators $D_i = 1\{T_i \leq C_i\}$ for each i .

If the number of observations M is large, it is efficient to use histograms. Form the histogram `obs` of all observed times \tilde{T}_i . That is, `obs[k]` counts the number of values \tilde{T}_i in the interval `(breaks[k], breaks[k+1])` for $k > 1$ and `[breaks[1], breaks[2]]` for $k = 1$. Also form the histogram `nco` of all uncensored times, i.e. those \tilde{T}_i such that $D_i = 1$. These two histograms are the arguments passed to `kaplan.meier`.

The vectors `km` and `lambda` returned by `kaplan.meier` are (histogram approximations to) the Kaplan-Meier estimator of $F(t)$ and its hazard rate $\lambda(t)$. Specifically, `km[k]` is an estimate of $F(\text{breaks}[k+1])$, and `lambda[k]` is an estimate of the average of $\lambda(t)$ over the interval $(\text{breaks}[k], \text{breaks}[k+1])$.

The histogram breaks must include 0. If the histogram breaks do not span the range of the observations, it is important to count how many survival times \bar{T}_i exceed the rightmost breakpoint, and give this as the value `upperobs`.

Value

A list with two elements:

<code>km</code>	Kaplan-Meier estimate of the survival time c.d.f. $F(t)$
<code>lambda</code>	corresponding Nelson-Aalen estimate of the hazard rate $\lambda(t)$

These are numeric vectors of length n .

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>
and Rolf Turner <rolfturner@posteo.net>

See Also

[reduced.sample](#), [km.rs](#)

<code>kernel.factor</code>	<i>Scale factor for density kernel</i>
----------------------------	--

Description

Returns a scale factor for the kernels used in density estimation for numerical data.

Usage

```
kernel.factor(kernel = "gaussian")
```

Arguments

<code>kernel</code>	String name of the kernel. Options are "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". (Partial matching is used).
---------------------	---

Details

Kernel estimation of a probability density in one dimension is performed by [density.default](#) using a kernel function selected from the list above.

This function computes a scale constant for the kernel. For the Gaussian kernel, this constant is equal to 1. Otherwise, the constant c is such that the kernel with standard deviation 1 is supported on the interval $[-c, c]$.

For more information about these kernels, see [density.default](#).

Value

A single number.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Martin Hazelton <Martin.Hazelton@otago.ac.nz>.

See Also

[density.default](#), [dkernel](#), [kernel.moment](#), [kernel.squint](#)

Examples

```
kernel.factor("rect")
# bandwidth for Epanechnikov kernel with half-width h=1
h <- 1
bw <- h/kernel.factor("epa")
```

kernel.moment

Incomplete Moment of Smoothing Kernel

Description

Computes the complete or incomplete m th moment of a smoothing kernel.

Usage

```
kernel.moment(m, r, kernel = "gaussian", mean=0, sd=1/kernel.factor(kernel))
```

Arguments

m	Exponent (order of moment). An integer.
r	Upper limit of integration for the incomplete moment. A numeric value or numeric vector. Set $r=Inf$ to obtain the complete moment.
kernel	String name of the kernel. Options are "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". (Partial matching is used).
mean, sd	Optional numerical values giving the mean and standard deviation of the kernel.

Details

Kernel estimation of a probability density in one dimension is performed by [density.default](#) using a kernel function selected from the list above. For more information about these kernels, see [density.default](#).

The function `kernel.moment` computes the integral

$$\int_{-\infty}^r t^m k(t) dt$$

where $k(t)$ is the selected kernel, r is the upper limit of integration, and m is the exponent or order.

Note that, if `mean` and `sd` are not specified, the calculations assume that $k(t)$ is the **standard form** of the kernel, which has support $[-1, 1]$ and standard deviation $sigma = 1/c$ where $c = kernel.factor(kernel)$.

The code uses the explicit analytic expressions when $m = 0, 1, 2$ and numerical integration otherwise.

Value

A single number, or a numeric vector of the same length as `r`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Martin Hazelton <Martin.Hazelton@otago.ac.nz>.

See Also

[density.default](#), [dkernel](#), [kernel.factor](#), [kernel.squint](#)

Examples

```
kernel.moment(1, 0.1, "epa")
curve(kernel.moment(2, x, "epa"), from=-1, to=1)
```

`kernel.squint`

Integral of Squared Kernel

Description

Computes the integral of the squared kernel, for the kernels used in density estimation for numerical data.

Usage

```
kernel.squint(kernel = "gaussian", bw=1)
```

Arguments

kernel	String name of the kernel. Options are "gaussian", "rectangular", "triangular", "epanechnikov", "biweight", "cosine" and "optcosine". (Partial matching is used).
bw	Bandwidth (standard deviation) of the kernel.

Details

Kernel estimation of a probability density in one dimension is performed by `density.default` using a kernel function selected from the list above.

This function computes the integral of the squared kernel,

$$R = \int_{-\infty}^{\infty} k(x)^2 dx$$

where $k(x)$ is the kernel with bandwidth `bw`.

Value

A single number.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Martin Hazelton <Martin.Hazelton@otago.ac.nz>.

See Also

`density.default`, `dkernel`, `kernel.moment`, `kernel.factor`

Examples

```
kernel.squint("gaussian", 3)

# integral of squared Epanechnikov kernel with half-width h=1
h <- 1
bw <- h/kernel.factor("epa")
kernel.squint("epa", bw)
```

Description

Compute the Kaplan-Meier and Reduced Sample estimators of a survival time distribution function, using histogram techniques

Usage

```
km.rs(o, cc, d, breaks)
```

Arguments

<code>o</code>	vector of observed survival times
<code>cc</code>	vector of censoring times
<code>d</code>	vector of non-censoring indicators
<code>breaks</code>	Vector of breakpoints to be used to form histograms.

Details

This function is needed mainly for internal use in **spatstat**, but may be useful in other applications where you want to form the Kaplan-Meier estimator from a huge dataset.

Suppose T_i are the survival times of individuals $i = 1, \dots, M$ with unknown distribution function $F(t)$ which we wish to estimate. Suppose these times are right-censored by random censoring times C_i . Thus the observations consist of right-censored survival times $\tilde{T}_i = \min(T_i, C_i)$ and non-censoring indicators $D_i = 1\{T_i \leq C_i\}$ for each i .

The arguments to this function are vectors `o`, `cc`, `d` of observed values of \tilde{T}_i , C_i and D_i respectively. The function computes histograms and forms the reduced-sample and Kaplan-Meier estimates of $F(t)$ by invoking the functions [kaplan.meier](#) and [reduced.sample](#). This is efficient if the lengths of `o`, `cc`, `d` (i.e. the number of observations) is large.

The vectors `km` and `hazard` returned by [kaplan.meier](#) are (histogram approximations to) the Kaplan-Meier estimator of $F(t)$ and its hazard rate $\lambda(t)$. Specifically, `km[k]` is an estimate of $F(\text{breaks}[k+1])$, and `lambda[k]` is an estimate of the average of $\lambda(t)$ over the interval $(\text{breaks}[k], \text{breaks}[k+1])$. This approximation is exact only if the survival times are discrete and the histogram breaks are fine enough to ensure that each interval $(\text{breaks}[k], \text{breaks}[k+1])$ contains only one possible value of the survival time.

The vector `rs` is the reduced-sample estimator, `rs[k]` being the reduced sample estimate of $F(\text{breaks}[k+1])$. This value is exact, i.e. the use of histograms does not introduce any approximation error in the reduced-sample estimator.

Value

A list with five elements

<code>rs</code>	Reduced-sample estimate of the survival time c.d.f. $F(t)$
<code>km</code>	Kaplan-Meier estimate of the survival time c.d.f. $F(t)$
<code>hazard</code>	corresponding Nelson-Aalen estimate of the hazard rate $\lambda(t)$
<code>r</code>	values of t for which $F(t)$ is estimated
<code>breaks</code>	the breakpoints vector

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>
and Rolf Turner <rolfturner@posteo.net>

See Also

[reduced.sample](#), [kaplan.meier](#)

knots.ewcdf*Jump Points of an Empirical Weighted Cumulative Distribution Function*

Description

Extract the knots (jump points) of an empirical cumulative distribution function.

Usage

```
## S3 method for class 'ewcdf'
knots(Fn, ...)
## S3 method for class 'ecdf'
knots(Fn, ...)
```

Arguments

Fn	An empirical cumulative distribution function (object of class "ecdf" or "ewcdf").
...	Ignored.

Details

The function **knots** is generic.

The function **knots.ecdf** is the method for the class "ecdf" of empirical cumulative distribution functions; objects of this class are created by **ecdf**.

The function **knots.ewcdf** is the method for the class "ewcdf" of empirical weighted cumulative distribution functions. Objects of class "ewcdf" are created by **ewcdf**.

The jump points (locations of increments) of the function Fn will be returned as a numeric vector.

Value

Numeric vector.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

ecdf, **ewcdf**, **quantile.ewcdf**

Examples

```
x <- c(1, 2, 5)
w <- runif(3)
e <- ewcdf(x,w)
knots(e)
```

mean.ewcdf*Mean of Empirical Cumulative Distribution Function*

Description

Calculates the mean of a (weighted or unweighted) empirical cumulative distribution function.

Usage

```
## S3 method for class 'ecdf'
mean(x, trim=0, ...)

## S3 method for class 'ewcdf'
mean(x, trim=0, ...)
```

Arguments

x An empirical cumulative distribution function (object of class "ecdf" created by [ecdf](#)) or a weighted empirical cumulative distribution function (object of class "ewcdf" created by [ewcdf](#)).

trim The fraction (0 to 0.5) of data values to be trimmed from each end of their range, before the mean is computed.

... Ignored.

Details

These functions are methods for the generic [mean](#) for the classes "ecdf" and "ewcdf".

They calculate the mean of the probability distribution corresponding to the cumulative distribution function **x**. This is equivalent to calculating the (weighted or unweighted) mean of the original data values.

For *weighted* empirical cumulative distribution functions (class "ewcdf") the weights will first be normalised so that they sum to 1. The result of `mean.ewcdf` is always an average or weighted average of the original data values. The argument **trim** is interpreted as a probability under this normalised distribution; the corresponding quantiles are computed, and data outside these quantiles is deleted before calculating the weighted mean.

Value

A single number.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

Generic [mean](#) and [weighted.mean](#).
[ecdf](#), [ewcdf](#) to create the cumulative distribution functions.
[stieltjes](#) for integration with respect to a cumulative distribution function.

Examples

```
x <- 1:5
mean(x)
mean(ecdf(x))
w <- 1:5
mean(ewcdf(x, w))
```

quantile.density *Quantiles of a Density Estimate*

Description

Given a kernel estimate of a probability density, compute quantiles.

Usage

```
## S3 method for class 'density'
quantile(x, probs = seq(0, 1, 0.25), names = TRUE,
         ..., warn = TRUE)
```

Arguments

<code>x</code>	Object of class "density" computed by a method for density
<code>probs</code>	Numeric vector of probabilities for which the quantiles are required.
<code>names</code>	Logical value indicating whether to attach names (based on <code>probs</code>) to the result.
<code>...</code>	Ignored.
<code>warn</code>	Logical value indicating whether to issue a warning if the density estimate <code>x</code> had to be renormalised because it was computed in a restricted interval.

Details

This function calculates quantiles of the probability distribution whose probability density has been estimated and stored in the object `x`. The object `x` must belong to the class "density", and would typically have been obtained from a call to the function [density](#).

The probability density is first normalised so that the total probability is equal to 1. A warning is issued if the density estimate was restricted to an interval (i.e. if `x` was created by a call to [density](#) which included either of the arguments `from` and `to`).

Next, the density estimate is numerically integrated to obtain an estimate of the cumulative distribution function $F(x)$. Then for each desired probability p , the algorithm finds the corresponding quantile q .

The quantile q corresponding to probability p satisfies $F(q) = p$ up to the resolution of the grid of values contained in x . The quantile is computed from the right, that is, q is the smallest available value of x such that $F(x) \geq p$.

Value

A numeric vector containing the quantiles.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

[quantile](#), [quantile.ewcdf](#), [CDF](#).

Examples

```
dd <- density(runif(10))
quantile(dd)
```

quantile.ewcdf

Quantiles of Weighted Empirical Cumulative Distribution Function

Description

Compute quantiles of a weighted empirical cumulative distribution function.

Usage

```
## S3 method for class 'ewcdf'
quantile(x, probs = seq(0, 1, 0.25),
          names = TRUE, ...,
          normalise = TRUE, type=1)
```

Arguments

x	A weighted empirical cumulative distribution function (object of class "ewcdf", produced by ewcdf) for which the quantiles are desired.
probs	probabilities for which the quantiles are desired. A numeric vector of values between 0 and 1.
names	Logical. If TRUE, the resulting vector of quantiles is annotated with names corresponding to probs.

...	Ignored.
normalise	Logical value indicating whether <i>x</i> should first be normalised so that it ranges between 0 and 1.
type	Integer specifying the type of quantile to be calculated, as explained in quantile.default . Only types 1, 2 and 4 are currently implemented.

Details

This is a method for the generic [quantile](#) function for the class *ewcdf* of empirical weighted cumulative distribution functions.

The quantile for a probability *p* is computed as the right-continuous inverse of the cumulative distribution function *x* (assuming *type*=1, the default).

If *normalise*=TRUE (the default), the weighted cumulative function *x* is first normalised to have total mass 1 so that it can be interpreted as a cumulative probability distribution function.

Value

Numeric vector of quantiles, of the same length as *probs*.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk> and Kevin Ummel.

See Also

[ewcdf](#), [quantile](#)

Examples

```
z <- rnorm(50)
w <- runif(50)
Fun <- ewcdf(z, w)
quantile(Fun, c(0.95,0.99))
```

Description

Return the inverse function of a cumulative distribution function.

Usage

```
quantilefun(x, ...)

## S3 method for class 'ecdf'
quantilefun(x, ..., type=1)

## S3 method for class 'ewcdf'
quantilefun(x, ..., type=1)
```

Arguments

x Data for which the quantile function should be calculated. Either an object containing data (such as a pixel image) or an object representing a cumulative distribution function (of class "ecdf" or "ewcdf").

... Other arguments passed to methods.

type Integer specifying the type of quantiles, as explained in [quantile.default](#). Only types 1, 2 and 4 are currently implemented.

Details

Whereas the command [quantile](#) calculates the quantiles of a dataset corresponding to desired probabilities p , the command [quantilefun](#) returns a function which can be used to compute any quantiles of the dataset.

If $f <- \text{quantilefun}(x)$ then f is a function such that $f(p)$ is the quantile associated with any given probability p . For example $f(0.5)$ is the median of the original data, and $f(0.99)$ is the 99th percentile of the original data.

If x is a pixel image (object of class "im") then the pixel values of x will be extracted and the quantile function of the pixel values is constructed.

If x is an object representing a cumulative distribution function (object of class "ecdf" or "ewcdf") then the quantile function of the original data is constructed.

Value

A function in the R language.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

[ewcdf](#), [quantile.ewcdf](#), [ecdf](#), [quantile](#)

Examples

```
## numeric data
z <- rnorm(50)
FZ <- ecdf(z)
QZ <- quantilefun(FZ)
QZ(0.5) # median value of z
if(interactive()) plot(QZ,xlim=c(0,1),xlab="probability",ylab="quantile of z")
```

reduced.sample

Reduced Sample Estimator using Histogram Data

Description

Compute the Reduced Sample estimator of a survival time distribution function, from histogram data

Usage

```
reduced.sample(nco, cen, ncc, show=FALSE, uppercen=0)
```

Arguments

nco	vector of counts giving the histogram of uncensored observations (those survival times that are less than or equal to the censoring time)
cen	vector of counts giving the histogram of censoring times
ncc	vector of counts giving the histogram of censoring times for the uncensored observations only
uppercen	number of censoring times greater than the rightmost histogram breakpoint (if there are any)
show	Logical value controlling the amount of detail returned by the function value (see below)

Details

This function is needed mainly for internal use in **spatstat**, but may be useful in other applications where you want to form the reduced sample estimator from a huge dataset.

Suppose T_i are the survival times of individuals $i = 1, \dots, M$ with unknown distribution function $F(t)$ which we wish to estimate. Suppose these times are right-censored by random censoring times C_i . Thus the observations consist of right-censored survival times $\tilde{T}_i = \min(T_i, C_i)$ and non-censoring indicators $D_i = 1\{T_i \leq C_i\}$ for each i .

If the number of observations M is large, it is efficient to use histograms. Form the histogram `cen` of all censoring times C_i . That is, `obs[k]` counts the number of values C_i in the interval `(breaks[k], breaks[k+1])` for $k > 1$ and `[breaks[1], breaks[2]]` for $k = 1$. Also form the histogram `nco` of all uncensored times, i.e. those \tilde{T}_i such that $D_i = 1$, and the histogram of all censoring times for which the survival time is uncensored, i.e. those C_i such that $D_i = 1$. These three histograms are the arguments passed to `kaplan.meier`.

The return value `rs` is the reduced-sample estimator of the distribution function $F(t)$. Specifically, `rs[k]` is the reduced sample estimate of $F(\text{breaks}[k+1])$. The value is exact, i.e. the use of histograms does not introduce any approximation error.

Note that, for the results to be valid, either the histogram breaks must span the censoring times, or the number of censoring times that do not fall in a histogram cell must have been counted in `uppercen`.

Value

If `show = FALSE`, a numeric vector giving the values of the reduced sample estimator. If `show=TRUE`, a list with three components which are vectors of equal length,

<code>rs</code>	Reduced sample estimate of the survival time c.d.f. $F(t)$
<code>numerator</code>	numerator of the reduced sample estimator
<code>denominator</code>	denominator of the reduced sample estimator

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>
and Rolf Turner <rolfturner@posteo.net>

See Also

[kaplan.meier](#), [km.rs](#)

rounding

Detect Numerical Rounding

Description

Given a numeric vector, determine whether the values have been rounded to a certain number of decimal places.

Usage

```
rounding(x)

## Default S3 method:
rounding(x)
```

Arguments

`x` A numeric vector, or an object containing numeric spatial coordinates.

Details

The function `rounding` is generic. Its purpose is to determine whether numerical values have been rounded to a certain number of decimal places.

The `spatstat` family of packages provides methods for rounding for various spatial objects.

For a numeric vector `x`, the default method `rounding.default` determines whether the values in `x` have been rounded to a certain number of decimal places.

- If the entries of `x` are not all integers, then `rounding(x)` returns the smallest number of digits `d` after the decimal point such that `round(x, digits=d)` is identical to `x`. For example if `rounding(x) = 2` then the entries of `x` are rounded to 2 decimal places, and are multiples of 0.01.
- If all the entries of `x` are integers, then `rounding(x)` returns `-d`, where `d` is the smallest number of digits *before* the decimal point such that `round(x, digits=-d)` is identical to `x`. For example if `rounding(x) = -3` then the entries of `x` are multiples of 1000. If `rounding(x) = 0` then the entries of `x` are integers but not multiples of 10.
- If all entries of `x` are equal to 0, a value of 0 is returned.

Value

An integer.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Rolf Turner <rolfturner@posteo.net>

See Also

`round.ppp` in package `spatstat.geom`.

Examples

```
rounding(c(0.1, 0.3, 1.2))
rounding(c(1940, 1880, 2010))
rounding(0)
```

Description

Computes the Stieltjes integral of a function f with respect to a function M .

Usage

`stieljes(f, M, ...)`

Arguments

<code>f</code>	The integrand. A function in the R language.
<code>M</code>	The cumulative function against which <code>f</code> will be integrated. An object of class "fv" or "stepfun".
...	Additional arguments passed to <code>f</code> .

Details

This command computes the Stieltjes integral

$$I = \int f(x)dM(x)$$

of a real-valued function $f(x)$ with respect to a nondecreasing function $M(x)$.

One common use of the Stieltjes integral is to find the mean value of a random variable from its cumulative distribution function $F(x)$. The mean value is the Stieltjes integral of $f(x) = x$ with respect to $F(x)$.

The argument `f` should be a function in the R language. It should accept a numeric vector argument `x` and should return a numeric vector of the same length.

The argument `M` should be either a step function (object of class "stepfun") or a function value table (object of class "fv"). Objects of class "stepfun" are returned by [ecdf](#), [ewcdf](#), and other utilities.

Value

A list containing the value of the Stieltjes integral computed using each of the versions of the function `M`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

Examples

```
x <- runif(100)
w <- runif(100)
H <- ewcdf(x, w)
stieltjes(function(x) { x^2 }, H)
```

 transformquantiles *Transform the Quantiles*

Description

Apply a transformation to the quantiles of a vector, or to the quantiles of the pixel values in a pixel image.

Usage

```
transformquantiles(X, uniform = FALSE, reverse = FALSE, ...)
```

Arguments

X	A numeric vector, matrix, array, or a pixel image (object of class "im").
uniform	Logical value specifying whether each quantile value should be replaced by the corresponding cumulative probability (called <i>histogram equalisation</i> , <i>transformation to uniformity</i> or <i>probability integral transformation</i>).
reverse	Logical value specifying whether to swap the upper and lower quantiles.
...	Ignored.

Details

The argument X may be a vector, matrix, array, or a pixel image (object of class "im").

The algorithm will first extract the entries or pixel values of X as a vector, and sort the values into ascending order.

If uniform=TRUE, the entries in this vector will be replaced by the corresponding cumulative probabilities (the kth smallest value will be replaced by the number $(k-0.5)/n$ where n is the total number of values).

If reverse=TRUE, the resulting vector will be reversed so that it is in descending order (so that the kth smallest value will be swapped with the kth largest value).

Finally the transformed values will be replaced into the original positions in the vector, matrix, array, or pixel image.

The case uniform=TRUE, reverse=FALSE is called *transformation to uniformity*, the *probability integral transformation*, *histogram equalisation*, or *quantile transformation*. The resulting values are uniformly distributed between 0 and 1; a histogram of the values in X is flat.

Value

Another object of the same type as X.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

To apply an arbitrary function f to the pixel values in an image, use the idiom $X[] \leftarrow f(X[])$.

Examples

```
X <- c(3, 5, 1, 2, 4)
transformquantiles(X, reverse=TRUE)
transformquantiles(X, uniform=TRUE)
transformquantiles(X, uniform=TRUE, reverse=TRUE)
```

uniquemap.default *Map Duplicate Entries to Unique Entries*

Description

Determine whether entries in a vector (or rows in a matrix or data frame) are duplicated, choose a unique representative for each set of duplicates, and map the duplicates to the unique representative.

Usage

```
uniquemap(x)

## Default S3 method:
uniquemap(x)

## S3 method for class 'data.frame'
uniquemap(x)

## S3 method for class 'matrix'
uniquemap(x)
```

Arguments

x A vector, data frame or matrix, or another type of data.

Details

The function `uniquemap` is generic, with methods for point patterns, data frames, and a default method.

The default method expects a vector. It determines whether any entries of the vector x are duplicated, and constructs a mapping of the indices of x so that all duplicates are mapped to a unique representative index.

The result is an integer vector u such that $u[j] = i$ if the entries $x[i]$ and $x[j]$ are identical and point i has been chosen as the unique representative. The entry $u[i] = i$ means either that point i is unique, or that it has been chosen as the unique representative of its equivalence class.

The method for `data.frame` determines whether any rows of the data frame `x` are duplicated, and constructs a mapping of the row indices so that all duplicate rows are mapped to a unique representative row.

Value

An integer vector.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>, Rolf Turner <rolfturner@posteo.net> and Ege Rubak <rubak@math.aau.dk>.

See Also

[duplicated](#).

`uniquemap.ppp` in **spatstat.geom**

Examples

```
x <- c(3, 5, 2, 4, 2, 3)
uniquemap(x)

df <- data.frame(A=x, B=42)
uniquemap(df)

z <- cbind(x, 10-x)
uniquemap(z)
```

unnormdensity

Weighted kernel smoother

Description

An unnormalised version of kernel density estimation where the weights are not required to sum to 1. The weights may be positive, negative or zero.

Usage

```
unnormdensity(x, ..., weights = NULL, defaults)
```

Arguments

<code>x</code>	Numeric vector of data
<code>...</code>	Optional arguments passed to density.default . Arguments must be <i>named</i> .
<code>*</code>	

weights	Optional numeric vector of weights for the data. The default is equivalent to assuming a weight of 1 for each observation.
defaults	Optional, named list of arguments passed to density.default . These will be overridden by arguments in

Details

This is an alternative to the standard R kernel density estimation function [density.default](#).

The standard [density.default](#) requires the weights to be nonnegative numbers that add up to 1, and returns a probability density (a function that integrates to 1).

This function `unnormdensity` does not impose any requirement on the `weights` except that they be finite. Individual weights may be positive, negative or zero. The result is a function that does not necessarily integrate to 1 and may be negative. The result is the convolution of the kernel k with the weighted data,

$$f(x) = \sum_i w_i k(x - x_i)$$

where x_i are the data points and w_i are the weights.

The argument `weights` should be a numeric vector of the same length as `x`, or a single numeric value. The default is to assume a weight of 1 for each observation in `x`.

The algorithm first selects the kernel bandwidth by applying [density.default](#) to the data `x` with normalised, positive weight vector `w = abs(weights)/sum(abs(weights))` and extracting the selected bandwidth. Then the result is computed by applying [density.default](#) to `x` twice using the normalised positive and negative parts of the weights.

Note that the arguments `...` must be passed by name, i.e. in the form `(name=value)`. Arguments that do not match an argument of [density.default](#) will be ignored *silently*.

Value

Object of class "density" as described in [density.default](#).

The result contains an additional component `$kernel` giving the name of the smoothing kernel that was used.

Warning

If `weights` is not specified, the default is to assign a weight $w_i = 1$ to each observation x_i .

This is not the same behaviour as in [density.default](#) which effectively assumes a weight of $1/n$ for each observation x_i where $n = \text{length}(x)$.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au> and Rolf Turner <rolfturner@posteo.net>

See Also

[density.default](#)

Examples

```
d <- unnormdensity(1:3, weights=c(-1,0,1), bw=0.3)
if(interactive()) plot(d)
```

weighted.median	<i>Weighted Median, Quantiles or Variance</i>
-----------------	---

Description

Compute the median, quantiles or variance of a set of numbers which have weights associated with them.

Usage

```
weighted.median(x, w, na.rm = TRUE, type=2, collapse=FALSE)
weighted.quantile(x, w, probs=seq(0,1,0.25), na.rm = TRUE, type=4, collapse=FALSE)
weighted.var(x, w, na.rm = TRUE)
```

Arguments

<code>x</code>	Data values. A vector of numeric values, for which the median or quantiles are required.
<code>w</code>	Weights. A vector of nonnegative numbers, of the same length as <code>x</code> . If <code>w</code> is missing or <code>NULL</code> , the default weights are all equal to 1.
<code>probs</code>	Probabilities for which the quantiles should be computed. A numeric vector of values between 0 and 1.
<code>na.rm</code>	Logical. Whether to ignore NA values.
<code>type</code>	Integer specifying the rule for calculating the median or quantile, corresponding to the rules available for quantile . The currently available choices are <code>type=1</code> , <code>2</code> , <code>3</code> and <code>4</code> . See Details.
<code>collapse</code>	Logical value specifying whether duplicated values in <code>x</code> should be pooled (replacing them by a unique <code>x</code> value whose weight is the sum of the associated weights).

Details

The i th observation $x[i]$ is treated as having a weight proportional to $w[i]$.

The weighted sample median is a value m such that the total weight of data less than or equal to m is equal to half the total weight. More generally, the weighted sample quantile with probability p is a value q such that the total weight of data less than or equal to q is equal to p times the total weight.

Define the weighted empirical cumulative distribution function

$$F(x) = \sum_{i:x_i \leq x} w_i / \sum_{i=1}^n w_i$$

That is, $F(x)$ is the fraction of total weight that is associated with data values x_i less than or equal to the value x .

The weighted quantile for probability p is a number q defined so that $F(q) = p$ wherever possible. There are different definitions of the quantile depending on how this should be achieved.

For unweighted data, there are 9 different definitions of the sample median and sample quantile, which enjoy slightly different properties. These 9 different definitions are explained in the help for `quantile.default`. The user's choice of algorithm is selected using the argument `type`.

For weighted data, the first 4 of the 9 definitions of sample quantile have been generalised to weighted quantiles. The functions `weighted.median` and `weighted.quantile` documented here provide these definitions of weighted sample quantile. The user's choice of algorithm is again selected using the argument `type`.

Suppose the data values have been arranged in increasing order as $x_{[1]} \leq x_{[2]} \leq \dots \leq x_{[n]}$. If one of the data values $x_{[k]}$ satisfies $F(x_{[k]}) = p$ exactly, then

- if `type=1`, `type=3` or `type=4`, the code returns this value, $x_{[k]}$;
- if `type=2`, the code returns the average of this value and the next largest value, $(x_{[k]} + x_{[k+1]})/2$.

If there is no data value satisfying $F(x_{[k]}) = p$ exactly, then the code finds the two adjacent values $x_{[k]}$ and $x_{[k+1]}$ which satisfy $F(x_{[k]}) < p$ and $F(x_{[k+1]}) > p$, and defines the quantile as follows:

- if `type=1` or `type=2`, the code returns the larger value $x_{[k+1]}$;
- if `type=3`, the code returns the value which minimises the discrepancy, that is, if we define $d_k = |F(x_{[k]}) - p|$ and $d_{k+1} = |F(x_{[k+1]}) - p|$, then
 - if $d_k < d_{k+1}$, the code returns $x_{[k]}$;
 - if $d_k > d_{k+1}$, the code returns $x_{[k+1]}$;
 - if $d_k = d_{k+1}$, then the even-numbered value is returned, that is, the code returns $x_{[k^*]}$ where k^* equals k if k is even, and equals $k + 1$ if $k + 1$ is even.
- if `type=4`, the code returns a value obtained by linearly interpolating between the two adjacent values $x_{[k]}$ and $x_{[k+1]}$, that is, it returns the value $(1 - a)x_{[k]} + ax_{[k+1]}$ where $a = (p - F(x_{[k]}))/(F(x_{[k+1]}) - F(x_{[k]}))$.

For very small probabilities $p < F(x_{[1]})$ the value $x_{[1]}$ is returned. For very large probabilities $p > F(x_{[n]})$ the value $x_{[n]}$ is returned.

Type 1 is the left-continuous quantile function.

Type 2 is consistent with the traditional definition of the sample median.

Types 1 and 3 always return a value selected from the input data `x`, while types 2 and 4 sometimes return values that are interpolated between the input data values.

Note that the default settings are different for `weighted.median` and `weighted.quantile`.

The implementation of type 3 is experimental and may be changed.

Value

`weighted.median` returns a numeric value. `weighted.quantile` returns a numeric vector of the same length as `probs`.

Author(s)

Adrian Baddeley <Adrian.Baddeley@curtin.edu.au>.

See Also

[quantile](#), [median](#).

Examples

```
x <- 1:20
w <- runif(20)
weighted.median(x, w)
weighted.quantile(x, w, probs=(0:4)/4)
weighted.var(x, w)
```

whist

Weighted Histogram

Description

Computes the weighted histogram of a set of observations with a given set of weights.

Usage

```
whist(x, breaks, weights = NULL, method=c("C", "interpreted"))
```

Arguments

<code>x</code>	Numeric vector of observed values.
<code>breaks</code>	Vector of breakpoints for the histogram.
<code>weights</code>	Numeric vector of weights for the observed values.
<code>method</code>	Developer use only. A character string specifying whether to use internal C code (<code>method="C"</code> , the default) or interpreted R code (<code>method="interpreted"</code>).

Details

This low-level function computes (but does not plot) the weighted histogram of a vector of observations `x` using a given vector of `weights`.

The arguments `x` and `weights` should be numeric vectors of equal length. They may include NA or infinite values.

The argument `breaks` should be a numeric vector whose entries are strictly increasing. These values define the boundaries between the successive histogram cells. The `breaks` *do not* have to span the range of the observations.

There are $N-1$ histogram cells, where $N = \text{length}(\text{breaks})$. An observation $x[i]$ falls in the j th cell if $\text{breaks}[j] \leq x[i] < \text{breaks}[j+1]$ (for $j < N-1$) or $\text{breaks}[j] \leq x[i] \leq \text{breaks}[j+1]$

(for $j = N-1$). The weighted histogram value $h[j]$ for the j th cell is the sum of `weights[i]` for all observations `x[i]` that fall in the cell.

Note that, in contrast to the function `hist`, the function `whist` does not require the breakpoints to span the range of the observations `x`. Values of `x` that fall outside the range of `breaks` are handled separately; their total weight is returned as an attribute of the histogram.

Value

A numeric vector of length $N-1$ containing the histogram values, where $N = \text{length}(\text{breaks})$.

The return value also has attributes "low" and "high" giving the total weight of all observations that are less than the lowest breakpoint, or greater than the highest breakpoint, respectively.

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Examples

```
x <- rnorm(100)
b <- seq(-1,1,length=21)
w <- runif(100)
whist(x,b,w)
```

Index

- * **Adaptive smoothing**
 - bw.abram, 5
- * **Bandwidth selection**
 - bw.abram, 5
 - bw.taylor, 9
- * **Histogram equalisation**
 - transformquantiles, 44
- * **Probability integral transformation**
 - transformquantiles, 44
- * **Quantile transformation**
 - transformquantiles, 44
- * **arith**
 - whist, 50
- * **distribution**
 - spatstat.univar-package, 2
- * **manip**
 - transformquantiles, 44
- * **math**
 - firstdigit, 22
 - hotrod, 24
 - indefinteg, 25
 - integral, 26
 - integral.density, 27
 - rounding, 41
 - stiltjes, 42
 - weighted.median, 48
- * **methods**
 - densityBC, 15
 - dkernel, 19
 - integral.density, 27
 - kernel.factor, 29
 - kernel.moment, 30
 - kernel.squint, 31
 - mean.ewcdf, 35
 - quantile.density, 36
 - uniquemap.default, 45
- * **nonparametric**
 - bw.abram, 5
 - bw.abram.default, 6
- bw.pow, 8
- CDF, 11
- densityAdaptiveKernel, 12
- densityAdaptiveKernel.default, 12
- densityBC, 15
- dkernel, 19
- ewcdf, 21
- integral.density, 27
- kaplan.meier, 28
- kernel.factor, 29
- kernel.moment, 30
- kernel.squint, 31
- km.rs, 32
- knots.ewcdf, 34
- quantile.density, 36
- quantile.ewcdf, 37
- quantilefun, 38
- reduced.sample, 40
- spatstat.univar-package, 2
- transformquantiles, 44
- * **package**
 - spatstat.univar-package, 2
- * **smooth**
 - bw.taylor, 9
 - densityBC, 15
 - dkernel, 19
 - kernel.factor, 29
 - kernel.moment, 30
 - kernel.squint, 31
 - spatstat.univar-package, 2
 - unnormdensity, 46
- * **spatial**
 - integral, 26
 - kaplan.meier, 28
 - km.rs, 32
 - mean.ewcdf, 35
 - quantile.ewcdf, 37
 - quantilefun, 38
 - reduced.sample, 40

stieljes, 42
transformquantiles, 44
uniquemap.default, 45

* univar
 bw.taylor, 9
 CDF, 11
 ewcdf, 21
 integral.density, 27
 knots.ewcdf, 34
 mean.ewcdf, 35
 quantile.density, 36
 transformquantiles, 44

 bw.abram, 5, 7–9, 12
 bw.abram.default, 3, 5, 6, 6, 13, 14
 bw.nrd, 10
 bw.nrd0, 6, 8, 9
 bw.pow, 8
 bw.taylor, 9

 CDF, 11, 37
 CDF.density, 3, 4, 28

 density, 11, 36
 density.default, 3, 4, 6, 7, 13, 15–21, 27, 28, 30–32, 46, 47

 densityAdaptiveKernel, 3, 12
 densityAdaptiveKernel.default, 3, 12, 18
 densityBC, 3, 15, 21
 dkernel, 4, 18, 19, 21, 30–32
 dkernelBC, 4, 20
 duplicated, 46

 ecdf, 22, 34–36, 39, 43
 ewcdf, 3, 21, 34–39, 43

 firstdigit, 5, 22

 hist, 51
 hotrod, 4, 24

 indefinteg, 5, 25
 integral, 26, 27
 integral.density, 5, 26, 27
 integrate, 25, 26

 kaplan.meier, 4, 28, 33, 41
 kernel.factor, 4, 20, 29, 31, 32
 kernel.moment, 4, 20, 30, 30, 32
 kernel.squint, 4, 20, 30, 31, 31

 km.rs, 29, 32, 41
 knots, 34
 knots.ecdf (knots.ewcdf), 34
 knots.ewcdf, 4, 34

 lastdigit, 5
 lastdigit(firstdigit), 22

 match.kernel, 20, 21
 mean, 22, 35, 36
 mean.ecdf (mean.ewcdf), 35
 mean.ewcdf, 3, 22, 35
 median, 50

 ndigits, 5
 ndigits(firstdigit), 22

 pkernel, 4
 pkernel(dkernel), 19
 plot, 22
 print, 22

 qkernel, 4
 qkernel(dkernel), 19
 quantile, 22, 37–39, 48, 50
 quantile.default, 38, 39, 49
 quantile.density, 4, 11, 28, 36
 quantile.ewcdf, 3, 4, 22, 34, 37, 37, 39
 quantilefun, 4, 38
 quantilefun.ewcdf, 4

 reduced.sample, 4, 29, 33, 40

 rkernel, 4
 rkernel(dkernel), 19
 round, 42
 rounding, 23, 41
 rounding.default, 5

 spatstat.univar
 (spatstat.univar-package), 2

 spatstat.univar-package, 2
 stieljes, 5, 22, 36, 42
 summary, 22

 transformquantiles, 4, 44

 uniquemap (uniquemap.default), 45
 uniquemap.default, 5, 45
 unnormdensity, 3, 27, 46

 weighted.mean, 36

weighted.median, [4](#), [48](#)
weighted.quantile, [4](#)
weighted.quantile (weighted.median), [48](#)
weighted.var (weighted.median), [48](#)
whist, [3](#), [50](#)